

# Solar Neutrinos and UC Irvine

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Physics and Astronomy

# Beta Decay of Neutrons, and Postulate of the Neutrino

- Neutrons live about 15 minutes, and then decay to a proton plus a positron plus missing energy.
- Wolfgang Pauli in 1930 postulated that the energy was carried away by a light neutral particle of spin  $\frac{1}{2}$ , which he named the neutrino.
- He won the Nobel Prize in 1945 for the Pauli Exclusion Principle, which states that no two electrons can occupy the same physical state.





I have done a terrible thing, I have postulated a particle that cannot be detected.

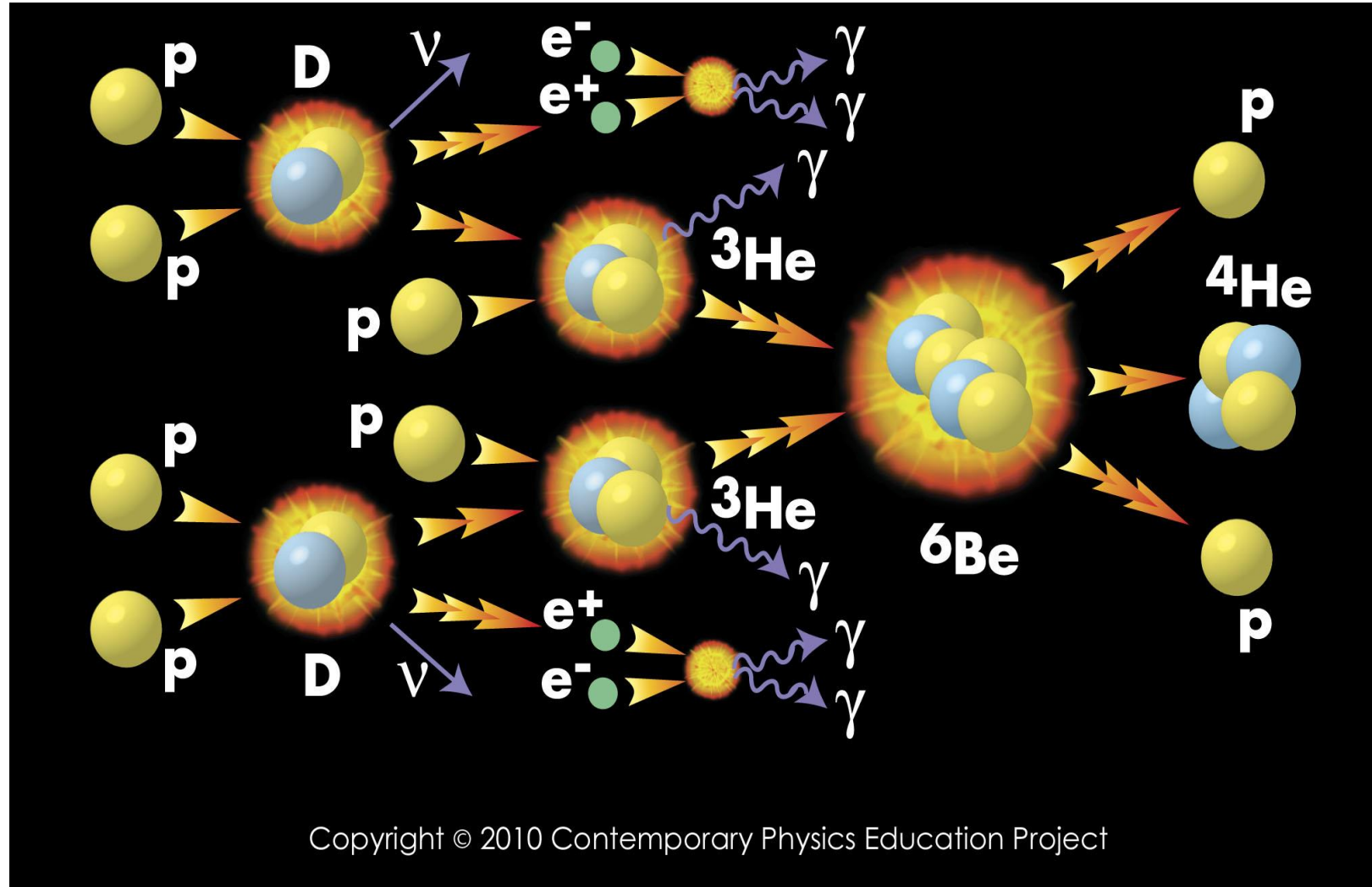
— *Wolfgang Pauli* —

AZ QUOTES

## Beta Decay Turned Around to Make Solar Deuterons

- Beta Decay:  $n \rightarrow p + e^- + \nu_e$  is a slow or weak interaction.
- Switch particle sides, making them their antiparticles:
- Proton to neutron conversion in a nucleus where neutron is better bound (inverse beta decay):  $p \rightarrow n + e^+ + \bar{\nu}_e$
- The start of solar fusion is then the merger of two protons, where one becomes converted to a neutron, and the proton and neutron combine to form a deuteron (pn):  $p + p \rightarrow d + e^+ + \bar{\nu}_e$
- Then a deuteron, which is larger than a proton and more strongly attracts another proton, can merge with another proton to form  $\text{He}^3$ .

Solar Fusion: Protons to deuterons to  $\text{He}^3$  to  $\text{He}^4$   
Two electron neutrinos produced



Cowan and Reines Detect Anti-neutrinos From Neutron Decays in Nuclear Reactors in 1956

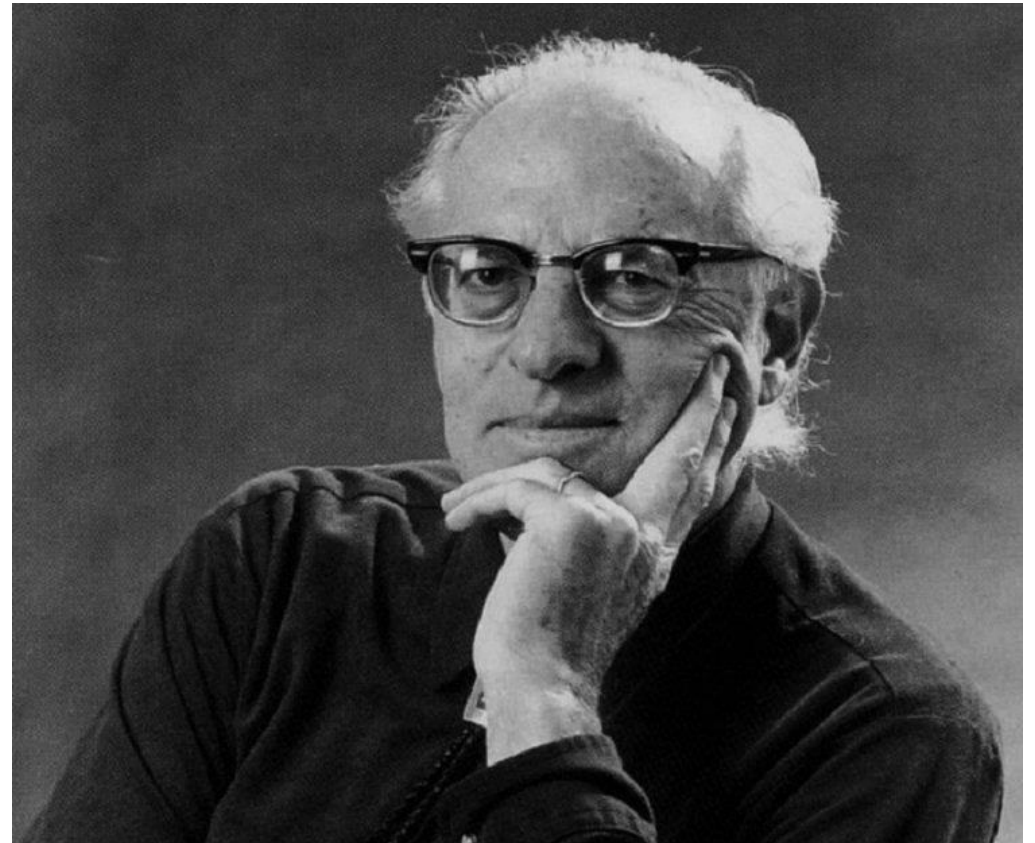
Fred Reines and Marty Perl share 1995 Nobel Prize in Physics

**Reines and Cowan**



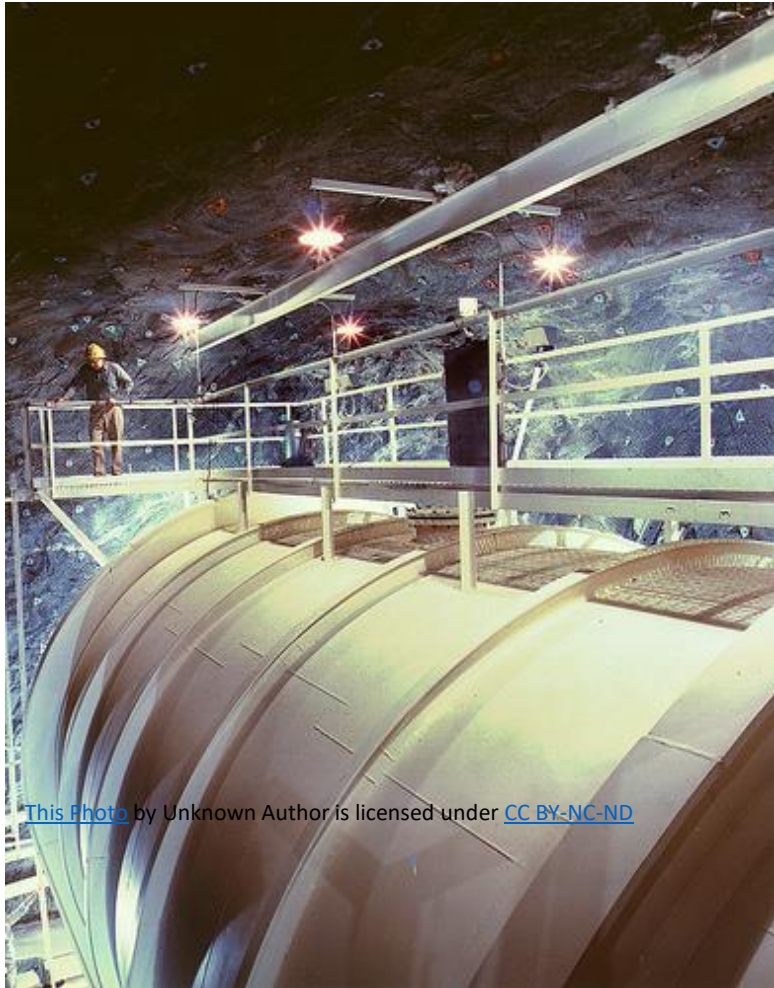
Frederick Reines (left) and Clyde L. Cowan, Jr. with the control equipment used in their first tentative observations of the neutrino at Hanford, Washington, in 1953. Their definitive detection of the (anti) neutrino was performed at Savannah River, Georgia, three years later. (Courtesy General Electric Co.)

**Fred Reines was Founding Dean of Physical Sciences at UC Irvine in 1965**





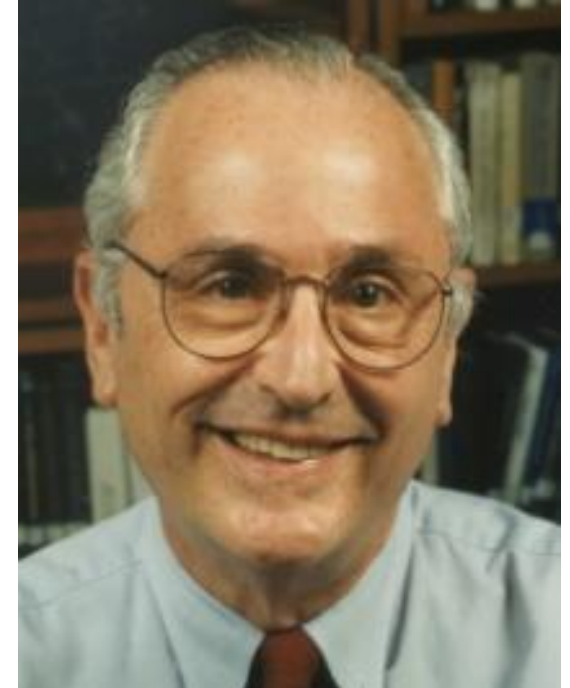
Ray Davis detects solar neutrinos with  $\nu_e + {}^{37}\text{Cl} \rightarrow {}^{37}\text{Ar} + e^-$   
John Bahcall did solar calculations. Only saw 1/3 of expected, due to neutrino oscillations.



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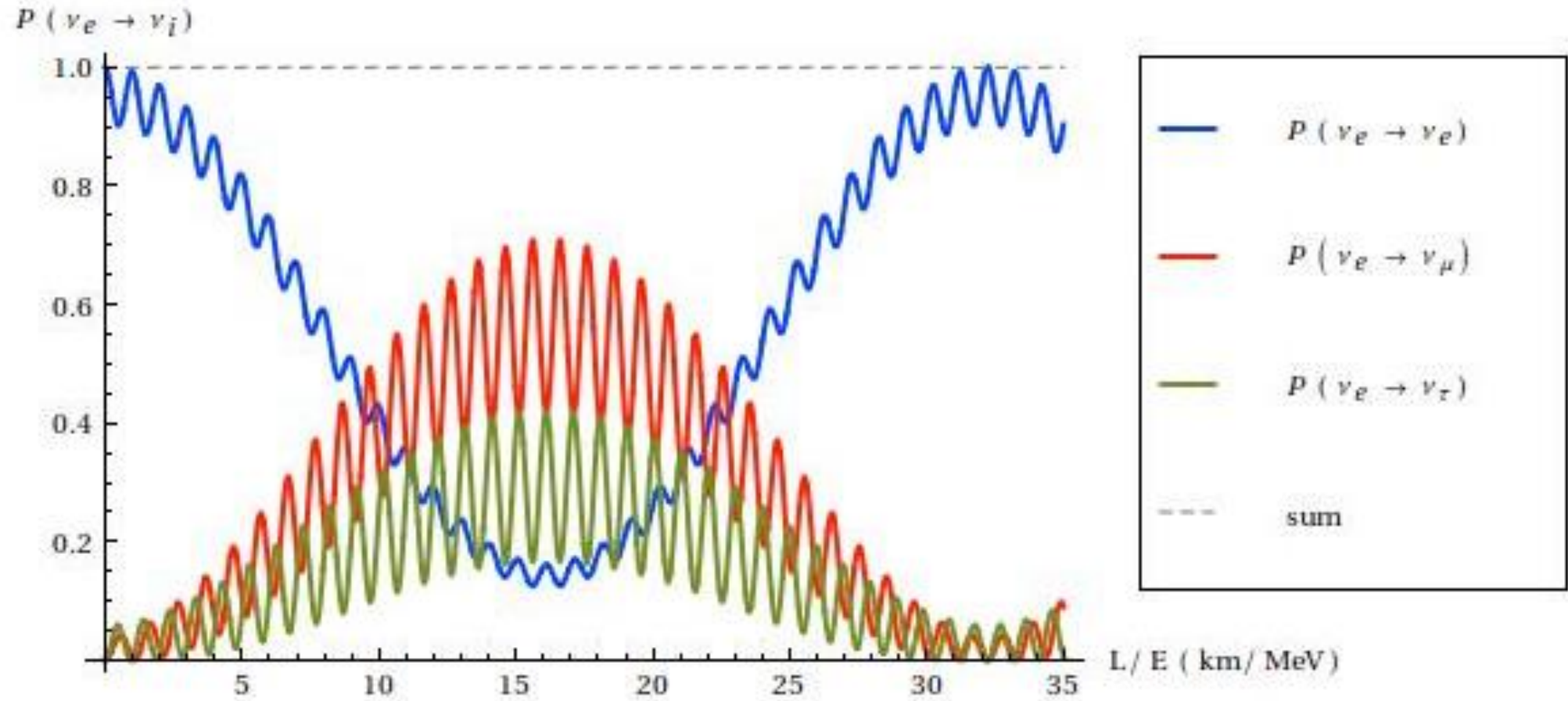
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## Two Other Leptons, Like the Electron, and Two Other Neutrinos

- Now have electron  $e$ , muon  $\mu$ , and tau  $\tau$ , all negatively charged.
- There are also their positively charged anti-particles with the same masses.
- Each one is co-produced with its own “flavor” of neutrino:
- $\nu_e$ ,  $\nu_\mu$ , and  $\nu_\tau$ , or their appropriate anti-neutrinos.
- These neutrinos don't have a definite mass, but are each mixtures of three fundamental neutrino states, each of which does have a definite mass:
- $\nu_1$ ,  $\nu_2$ , and, of course,  $\nu_3$ . They propagate as waves as different frequencies.
- These are named in the order of their increasing, but unknown, masses.
- Their masses are less than the incredibly light 1 eV. (The electron mass is 0.5 MeV)



Oscillation of the solar  $\nu_e$  (blue) to  $\nu_\mu$  (red), and  $\nu_\tau$  (green)



Sudbury Neutrino Observatory, SNO, counts  $\nu_e$ , and then all  $\nu$ 's

### NEUTRINOS FROM THE SUN

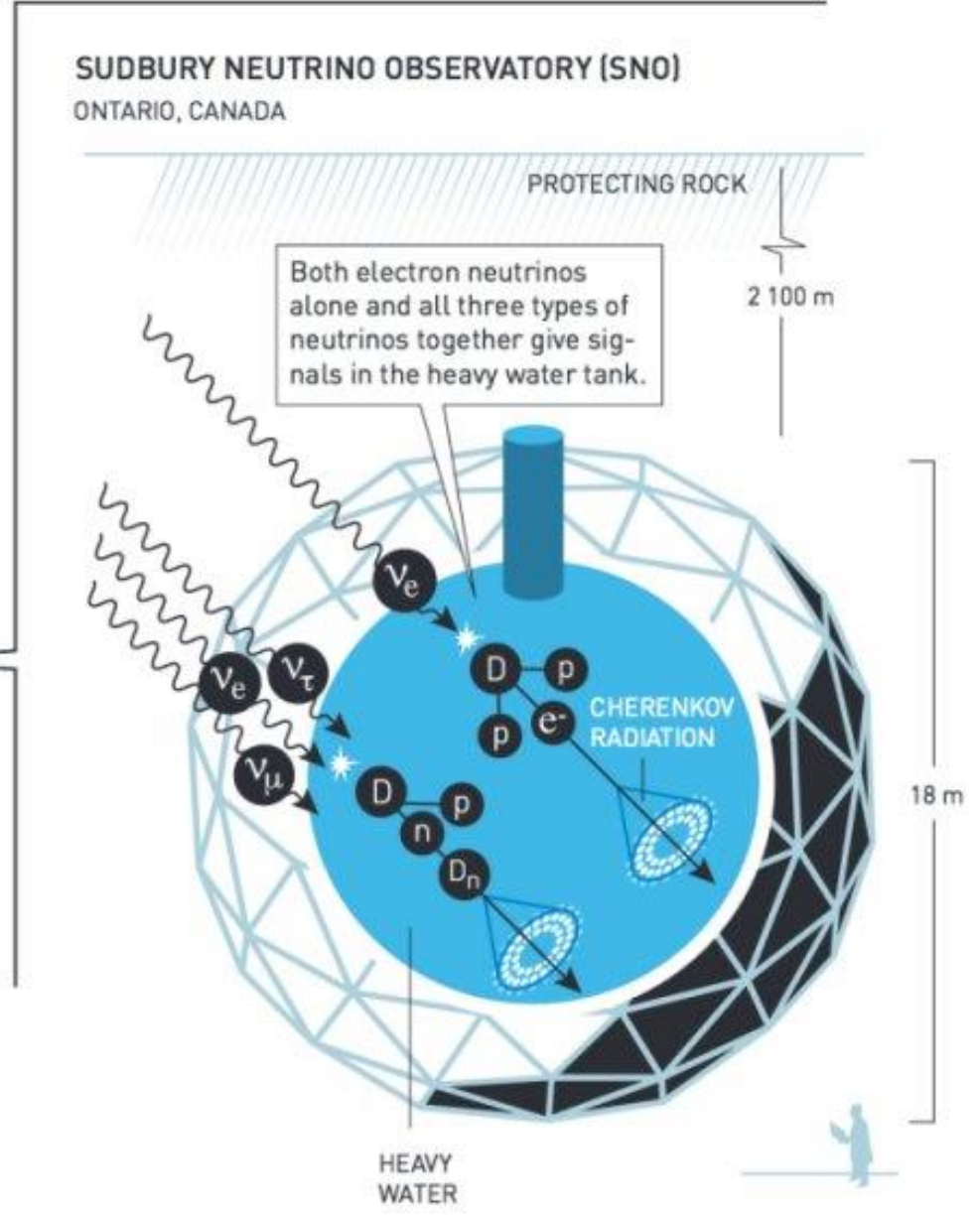
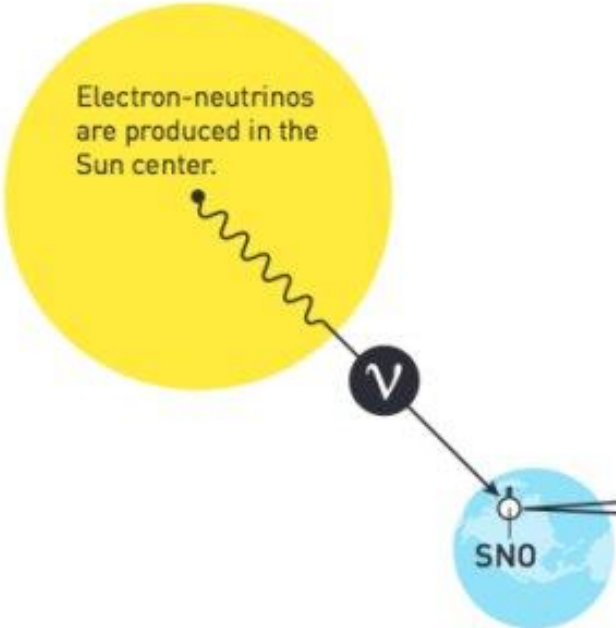
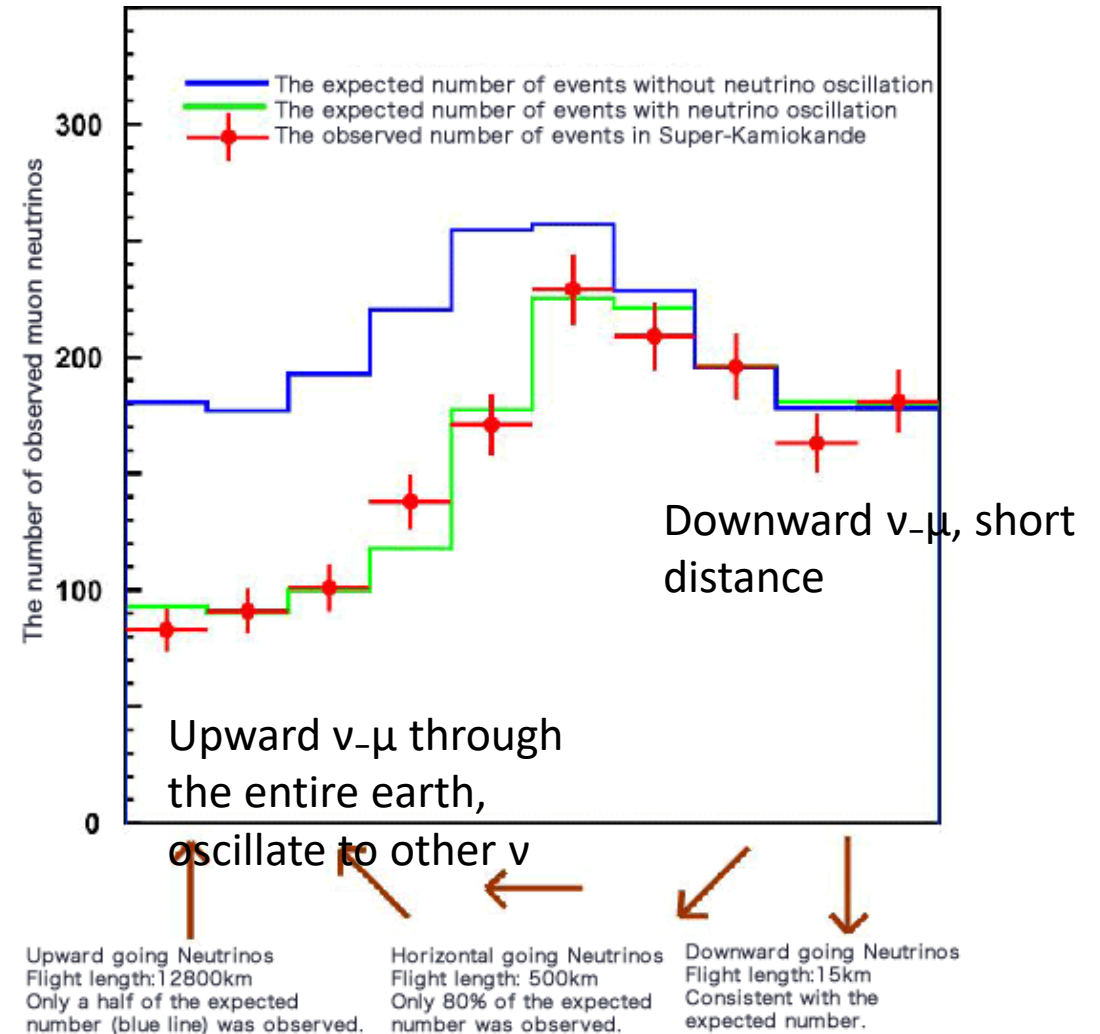


Illustration: © Johan Jarnestad/The Royal Swedish Academy of Sciences

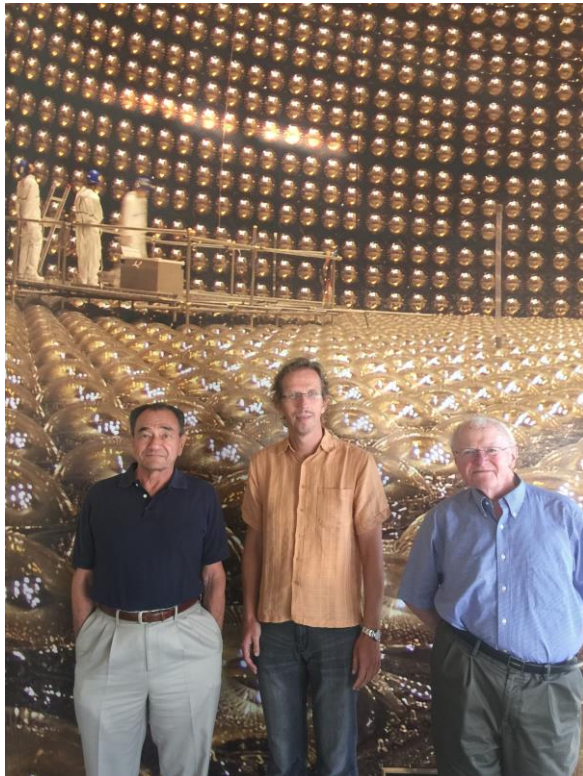
# Super-Kamiokande, Oscillation of atmospheric muon neutrinos traveling through the earth

- Cosmic rays strike the atmosphere and produce positive pions
- They decay to muons and  $\nu_{\mu}$ 's
- Traveling all the way through the earth, half of  $\nu_{\mu}$ 's oscillate to other neutrinos

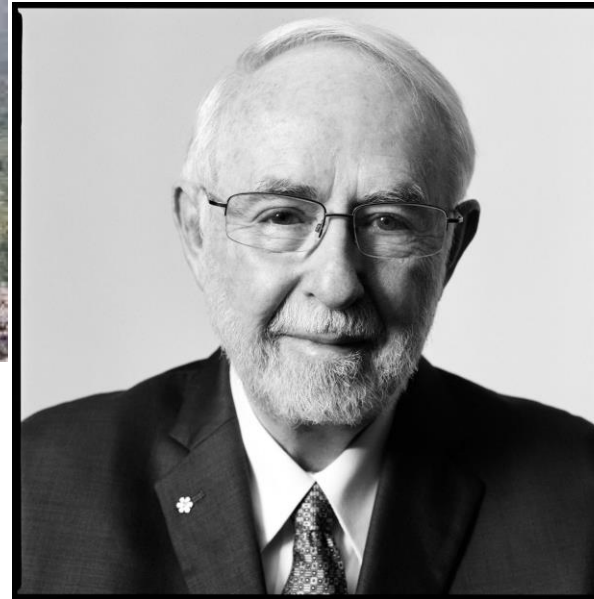


SNO lab Art McDonald and Super-K Takaaki Kajita  
Share 2015 Nobel Prize. Herb Chen, UCI, originator of SNO lab.

**UCI Super-K Hank Sobel, Michael Smy, Bill  
Kropp. SNO lab originator, Herb Chen**



**SNO lab director Art McDonald  
Super-K director Takaaki Kajita**





## 2002 Nobel Prize in Physics

- Masatoshi Koshiba, Kamiokande
- Supernova 1987A
- Raymond Davis, Jr.  
Solar Neutrinos

### Koshiba and Davis (2002)

- Nobel Prize for the detection of cosmic neutrinos



Koshiba



Davis

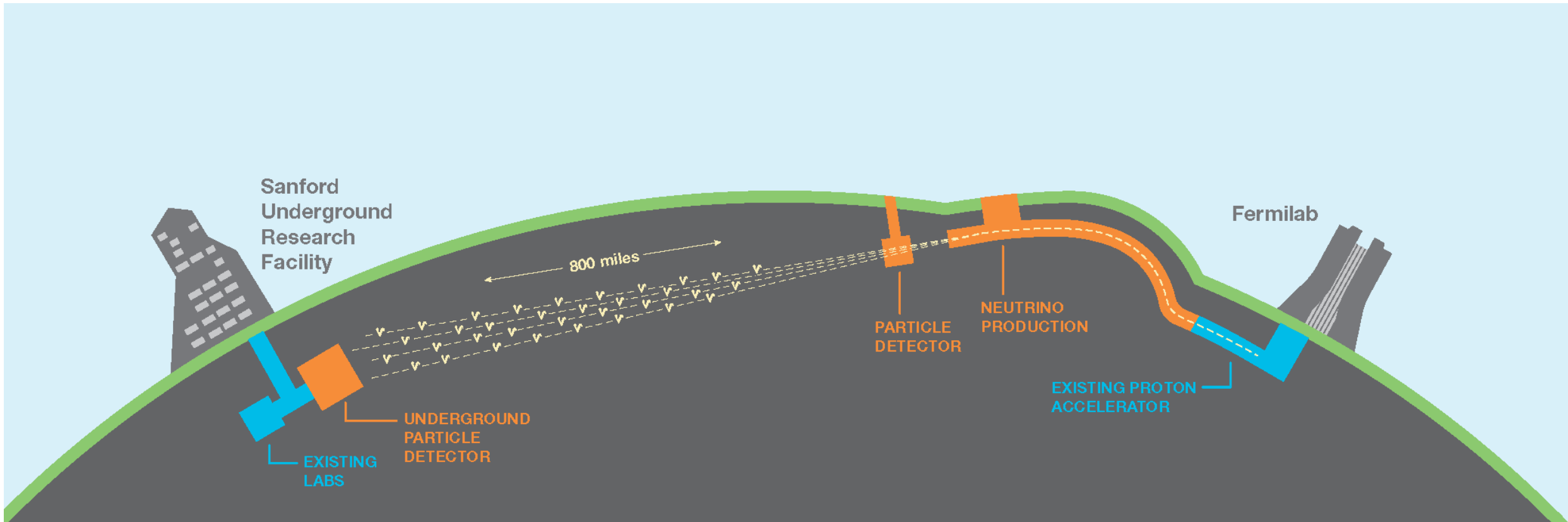
The 2016 Breakthrough Prize in Fundamental Physics was awarded to 5 neutrino oscillation experiments, including Super-K and SNO lab.

## DUNE

- The future US high energy physics facilities are led by the neutrino oscillation experiment DUNE in 2026
- 40 kilotons of liquid Argon target
- $\nu_{\mu}$  oscillating to  $\nu_e$
- And their anti-neutrino oscillation, to see what the differences between neutrinos and anti-neutrinos oscillations are.



# DUNE: 800 Mile Oscillation Path



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