#### The Noise Collector for Sparse Recovery in High Dimensions

#### Chrysoula Tsogka

Collaborators: M. Moscoso (Univ. Carlos III de Madrid), A. Novikov (Penn State),

G. Papanicolaou (Stanford University)

Department of Applied Mathematics, University of California, Merced Support: AFOSR FA9550-17-1-0238, AFOSR FA9550-18-1-0519.

・ロト ・日本・ ・ ヨト

## Table of contents

1 Inverse problems in wave propagation

#### 2 Applications

3 On  $\ell_1$  and (super)resolution

4 Uncertainty in the data

#### Conclusion

・ロト ・回ト ・ヨト ・



- Inverse problems aim to reconstruct a medium characteristics from knowledge of the response of the medium to a known incident field.
- In the context of the wave equation we seek to reconstruct the reflectivity by recording the medium's response to one or more known excitations.



- Inverse problems aim to reconstruct a medium characteristics from knowledge of the response of the medium to a known incident field.
- In the context of the wave equation we seek to reconstruct the reflectivity by recording the medium's response to one or more known excitations.



- Inverse problems aim to reconstruct a medium characteristics from knowledge of the response of the medium to a known incident field.
- In the context of the wave equation we seek to reconstruct the reflectivity by recording the medium's response to one or more known excitations.



- Inverse problems aim to reconstruct a medium characteristics from knowledge of the response of the medium to a known incident field.
- In the context of the wave equation we seek to reconstruct the reflectivity by recording the medium's response to one or more known excitations.
- We consider sparse reflectivities : often true in applications when the object to be imaged occupies a small part of the imaging scene



- Inverse problems aim to reconstruct a medium characteristics from knowledge of the response of the medium to a known incident field.
- In the context of the wave equation we seek to reconstruct the reflectivity by recording the medium's response to one or more known excitations.
- We consider sparse reflectivities : often true in applications when the object to be imaged occupies a small part of the imaging scene
- Study how uncertainty in the data affects the solution

### Table of contents

Inverse problems in wave propagation

#### 2 Applications

3 On  $\ell_1$  and (super)resolution

4 Uncertainty in the data

#### Conclusion

・ロト ・回ト ・ヨト ・

#### Radar



Here we want to retrieve the reflectivity on the ground. The measurements are obtained by sending pulses and collecting the echoes. The sensing matrix models how waves propagate from the airplane to the ground reflectors and back to the plane. The start-stop approximation is assumed. Uncertainty in the data : measurement noise, measurement location

#### Non-destructive testing



Uncertainty : unknown micro-structure of material ; concrete sample



< 🗇 >

Medical Ultrasound



Uncertainty : measurement noise, low SNR

・ロト ・回ト ・ヨト

# The imaging problem Passive case

• We consider here the passive imaging problem, where we seek to reconstruct the positions  $y_j$  and complex-valued amplitudes  $\rho_j$ , j = 1, ..., M of sources.



• There are N receivers at the array and K points in the image window IW. We assume that  $N \ll K$  and that the unknown  $\rho$  has sparse support  $M = \text{supp}(\rho)$ , M < N and typically  $M \ll K$ .

< □ > < 同 >

# The imaging problem Passive case

• We consider here the passive imaging problem, where we seek to reconstruct the positions  $y_j$  and complex-valued amplitudes  $\rho_j$ , j = 1, ..., M of sources.



ullet The imaging problem consists in solving a linear system of the form  $\mathcal{A} {\pmb \rho} = {\pmb b}$ 

with b the data, ho the unknown, and  $\mathcal A$  the sensing matrix

# The imaging problem

• We consider here the passive imaging problem, where we seek to reconstruct the positions  $y_j$  and complex-valued amplitudes  $\rho_j$ , j = 1, ..., M of sources.



# The imaging problem

• We consider here the passive imaging problem, where we seek to reconstruct the positions  $y_j$  and complex-valued amplitudes  $\rho_j$ , j = 1, ..., M of sources.



• The source imaging problem is considered here for simplicity. The active array imaging problem can be cast under the same linear algebra framework (i.e., solving  $A\rho = b$ ) even when multiple scattering is not negligible.

## Table of contents

1 Inverse problems in wave propagation

#### 2 Applications

#### (3) On $\ell_1$ and (super)resolution

#### 4 Uncertainty in the data

#### 5 Conclusion

・ロト ・回ト ・ヨト ・

#### Compressive sensing

Our problem consist in finding the solution  $\boldsymbol{\rho} \in \mathbb{C}^K$  of

 $\mathcal{A} \boldsymbol{\rho} = \boldsymbol{b}_{\delta},$ 

from highly incomplete ( $1 \ll N < K$ ) measurement data  $\boldsymbol{b}_{\delta} \in \mathbb{C}^N$ 

 $\boldsymbol{b}_{\delta} = \boldsymbol{b} + \boldsymbol{\delta} \boldsymbol{b},$ 

corrupted by the noise vector  $\delta b$ .

There exist infinitely many solutions to our problem, and thus, it is a priori not possible to find the correct one without some additional information.

The sparsity of  $\rho$  changes the imaging problem substantially because we can formulate it as an optimization problem which seeks the sparsest vector in  $\mathbb{C}^K$  that equates model and data.

#### $\rho^{\ell_1}: \min ||\rho||_1, \text{ subject to } \mathcal{A}\rho = b$

Theorem<sup>1</sup> : In the noiseless case the  $\ell_1$  approach gives the exact solution ! under the assumption (mutual coherence) :

 $\max_{k \neq k'} |\langle \boldsymbol{a}_k, \boldsymbol{a}_{k'} \rangle| \le 1/(2M), \quad \forall k, k' = 1, \dots, K.$ 

Here  $a_k$  : normalized columns of  $\mathcal{A}$ , i.e.,  $\|a_k\|_2 = 1, \ \forall k$ 

Theorem <sup>2</sup> : The same result can be obtained assuming the sensing matrix  $\mathcal{A}$  obeys the M-restricted isometry property which basically states that all sets of M-columns of  $\mathcal{A}$  behave approximately as an orthonormal system.

In both cases the assumptions translate to conditions on the discretization of IW since in imaging  $a_k = g(\boldsymbol{y}_k) / \|g(\boldsymbol{y}_k)\|_2$  with  $g(\boldsymbol{y}_k, \omega_l) = (G(\boldsymbol{x}_1, \boldsymbol{y}_k), G(\boldsymbol{x}_2, \boldsymbol{y}_k), \dots, G(\boldsymbol{x}_N, \boldsymbol{y}_k))^T$ 

1. Donoho & Elad '03

(日) (四) (日) (日) (日)

<sup>2.</sup> Candès & Tao '05

• A pair of columns  $a_i$  and  $a_j$  are coherent, if the corresponding grid-points are close to each other. Therefore the incoherence conditions can be satisfied only for coarse image discretizations that imply poor resolution.

メロト メタト メヨト メ

- A pair of columns  $a_i$  and  $a_j$  are coherent, if the corresponding grid-points are close to each other. Therefore the incoherence conditions can be satisfied only for coarse image discretizations that imply poor resolution.
- $\bullet\,$  To achieve high resolution we propose to modify the classical theory so as to allow for some coherence in  ${\cal A}.$

・ロト ・日本・ ・ ヨト

- A pair of columns  $a_i$  and  $a_j$  are coherent, if the corresponding grid-points are close to each other. Therefore the incoherence conditions can be satisfied only for coarse image discretizations that imply poor resolution.
- To achieve high resolution we propose to modify the classical theory so as to allow for some coherence in  $\mathcal{A}.$
- Let  $\rho \in \mathbb{C}^K$  be an *M*-sparse solution of  $\mathcal{A}\rho = \mathbf{b}$ . Define the index of its support *T*. For any  $j \in T$  define the corresponding vicinity of  $\mathbf{a}_j$  as

$$S_j = \left\{ k \text{ s.t. } |\langle \boldsymbol{a}_k, \boldsymbol{a}_j \rangle| \geqslant \frac{1}{3M} 
ight\}.$$

- A pair of columns  $a_i$  and  $a_j$  are coherent, if the corresponding grid-points are close to each other. Therefore the incoherence conditions can be satisfied only for coarse image discretizations that imply poor resolution.
- $\bullet\,$  To achieve high resolution we propose to modify the classical theory so as to allow for some coherence in  ${\cal A}.$
- Let  $\rho \in \mathbb{C}^K$  be an *M*-sparse solution of  $\mathcal{A}\rho = \mathbf{b}$ . Define the index of its support *T*. For any  $j \in T$  define the corresponding vicinity of  $\mathbf{a}_j$  as

$$S_j = \left\{ k \text{ s.t. } |\langle \boldsymbol{a}_k, \boldsymbol{a}_j \rangle| \geqslant rac{1}{3M} 
ight\}.$$

• For any vector  $\boldsymbol{\xi} \in \mathbb{C}^K$  define its coherent misfit to  $\boldsymbol{
ho}$  as

$$\mathbf{Co}(\boldsymbol{\rho}, \boldsymbol{\xi}) = \sum_{j \in T} \left| \rho_j - \sum_{k \in S_j} \langle \boldsymbol{a}_j, \boldsymbol{a}_k \rangle \xi_k \right|,$$

and its incoherent remainder

$$\mathbf{In}(\boldsymbol{\rho},\boldsymbol{\xi}) = \sum_{k \notin \Upsilon} |\xi_k|, \Upsilon = \cup_{j \in T} S_j.$$

A D F A A F F A

#### Our abstract result

#### Theorem

Let  $\rho$  be an *M*-sparse solution of  $A\rho = \mathbf{b}$ , and suppose vicinities  $S_j$  do not overlap. Let  $\rho_{\delta}$  be the minimal  $\ell_1$ -norm solution of the noisy problem

 $\min \|\boldsymbol{\rho}_{\delta}\|_{\ell_1}, \text{ subject to } \mathcal{A}\boldsymbol{\rho}_{\delta} = \boldsymbol{b}_{\delta},$ 

with  $\|\boldsymbol{b} - \boldsymbol{b}_{\delta}\|_{\ell_2} \leq \delta$ . Then

 $\mathbf{Co}(\boldsymbol{\rho}, \boldsymbol{\rho}_{\delta}) \leqslant 3\gamma\delta,$ 

and

 $\mathbf{In}(\boldsymbol{\rho}, \boldsymbol{\rho}_{\delta}) \leqslant 5\gamma\delta.$ 

If the noise level  $\delta = 0$ , and the number of vectors in  $\Upsilon = \bigcup_{j \in T} S_j$  is smaller than the rank of A, we have exact recovery :  $\rho_{\delta} = \rho$ .

• This theorem says that when the vicinities associated to different sources do not overlap,  $\ell_1$  minimization determines their location with high precision.

(日) (日) (日) (日)

- This theorem says that when the vicinities associated to different sources do not overlap,  $\ell_1$  minimization determines their location with high precision.
- Allowing for the columns of A to be coherent inside the vicinities is crucial as it permits for a fine discretization to be used which in turn implies high resolution imaging.

- This theorem says that when the vicinities associated to different sources do not overlap,  $\ell_1$  minimization determines their location with high precision.
- Allowing for the columns of A to be coherent inside the vicinities is crucial as it permits for a fine discretization to be used which in turn implies high resolution imaging.
- When there is noise in the data the location of the sources is not retrieved exactly but both the coherent and the incoherent misfit are small.

< □ > < 同 >

- This theorem says that when the vicinities associated to different sources do not overlap,  $\ell_1$  minimization determines their location with high precision.
- Allowing for the columns of A to be coherent inside the vicinities is crucial as it permits for a fine discretization to be used which in turn implies high resolution imaging.
- When there is noise in the data the location of the sources is not retrieved exactly but both the coherent and the incoherent misfit are small.
- The incoherent misfit concerns the support of  $\rho_{\delta}$  far from the true sources location, *i.e.*, outside the vicinities. This type of noise in the reconstruction image is usually referred to as grass.

- This theorem says that when the vicinities associated to different sources do not overlap,  $\ell_1$  minimization determines their location with high precision.
- Allowing for the columns of A to be coherent inside the vicinities is crucial as it permits for a fine discretization to be used which in turn implies high resolution imaging.
- When there is noise in the data the location of the sources is not retrieved exactly but both the coherent and the incoherent misfit are small.
- The incoherent misfit concerns the support of  $\rho_{\delta}$  far from the true sources location, *i.e.*, outside the vicinities. This type of noise in the reconstruction image is usually referred to as grass.
- The coherent misfit concerns the approximation of ρ by degrees of freedom inside the vicinities. The coherent misfit being small means that ρ can be well approximated by a linear combination of the values of ρ<sub>δ</sub> inside the vicinities of the true support.

(日) (同) (三) (三)

## $\ell_1$ and super-resolution

#### • Exact recovery in the noiseless case



Vicinities overlap

## $\ell_1$ and super-resolution

#### • Exact recovery in the noiseless case



Vicinities overlap

• Robustness to noise (under smallness assumption) : the solution is decomposed in two parts, the coherent part,  $\rho_c$ , which is supported in the vicinities  $S_j$  and is close to the true  $\rho$  and the incoherent part,  $\rho_i$ , usually referred to as grass. The grass is supported away from the vicinities  $S_j$  and is shown to be small.

### $\ell_1$ and super-resolution

- $\bullet$  Other works where super-resolution has been studied for highly coherent model matrices  ${\cal A}$  that arise in imaging
  - A. Fannjiang and W. Liao, *Coherence pattern-guided compressive sensing with unresolved grids*, SIAM J. Imaging Sci. 5 (2012), pp. 179–202.
  - L. Borcea and I. Kocyigit, *Resolution analysis of imaging with*  $\ell_1$  *optimization*, SIAM J. Imaging Sci. 8 (2015), pp. 3015–3050.
  - L. Borcea and I. Kocyigit, A multiple measurement vector approach to synthetic aperture radar imaging, SIAM J. Imaging Sci. 11 (2018), pp. 770–801.

These works include results regarding the robustness of super-resolution in the presence of noise.

(日) (同) (三) (三)

## Table of contents

1 Inverse problems in wave propagation

2 Applications

3 On  $\ell_1$  and (super)resolution

Uncertainty in the data

#### 5 Conclusion

・ロト ・回ト ・ヨト

### What about uncertainty in the data (noise)

Our theory implies that a key to control the noise is the constant  $\gamma$ ,

$$\gamma = \sup_{\boldsymbol{b}} \frac{\|\boldsymbol{\rho}\|_{\ell_1}}{\|\boldsymbol{b}\|_{\ell_2}}$$
 where  $\boldsymbol{\rho}$  is the minimal  $\ell_1$  – solution of  $\mathcal{A}\boldsymbol{\rho} = \boldsymbol{b}$ .

For the imaging problem we have  $\gamma = O(\sqrt{N})$  which is not satisfactory.



N = 625 measurements. K = 1681 pixels in the images. 100% noise.

### What about uncertainty in the data (noise)

Our theory implies that a key to control the noise is the constant  $\gamma$ ,

$$\gamma = \sup_{\boldsymbol{b}} \frac{\|\boldsymbol{\rho}\|_{\ell_1}}{\|\boldsymbol{b}\|_{\ell_2}}$$
 where  $\boldsymbol{\rho}$  is the minimal  $\ell_1$  – solution of  $\mathcal{A}\boldsymbol{\rho} = \boldsymbol{b}$ .

For the imaging problem we have  $\gamma = O(\sqrt{N})$  which is not satisfactory.



N = 1369 measurements. K = 1681 pixels in the images. 100% noise.

#### What about uncertainty in the data (noise)

The most commonly used approach to deal with noise is to solve the  $\ell_2\text{-relaxed}$  form of the optimization problem

$$\boldsymbol{\rho}_{\lambda} = \arg\min_{\boldsymbol{\rho}} \left( \lambda \|\boldsymbol{\rho}\|_{\ell_1} + \frac{1}{2} \|\boldsymbol{\mathcal{A}}\boldsymbol{\rho} - \boldsymbol{b}\|_{\ell_2}^2 \right), \tag{1}$$

known as Lasso in the statistics literature [R. Tibshirani '96, Chen & D.Donoho '94, F.Santosa & W.Symes '86].

There are sufficient conditions, depending on the SNR and the value of  $\lambda$  (tuning parameter which should be adequately chosen), for the support of  $\rho_{\lambda}$  to be contained within the true support.

### Tuning $\lambda$ in Lasso



▷ Develop a method for exact support recovery in the presence of noise. No tuning parameters. No a priory knowledge on the level of noise required.

< □ > < 同 >

#### The Noise Collector

We propose to solve instead

$$( oldsymbol{
ho}_{ au}, oldsymbol{\eta}_{ au}) = rg \min_{oldsymbol{
ho}, oldsymbol{\eta}} ( au \| oldsymbol{
ho} \|_{l_1} + \| oldsymbol{\eta} \|_{l_1}),$$
  
subject to  $\mathcal{A} oldsymbol{
ho} + \mathcal{C} oldsymbol{\eta} = oldsymbol{b} + \delta oldsymbol{b},$ 

where C is the *Noise Collector* matrix  $C \in \mathbb{C}^{N \times \Sigma}$ ,  $\Sigma \gg K$  and  $\tau$  is an O(1) weight that is independent of the dimension of the problem and the level of noise.

This minimization problem can be understood as a relaxation of

$$oldsymbol{
ho}_* = rg\min_{oldsymbol{
ho}} \|oldsymbol{
ho}\|_{\ell_1}, ext{ subject to } \mathcal{A} \, oldsymbol{
ho} = oldsymbol{b} + \delta oldsymbol{b},$$

It works by absorbing *all* the noise, and possibly some signal, in  $\mathcal{C}\eta_{\tau}$ .

 $\eta$  does not correspond to a physical quantity. It is introduced to provide a fictitious source distribution (an appropriate linear combination of the columns of C) that produces a good approximation to  $\delta b$ .

### Construction of the Noise Collector

- (i) Columns of C should be sufficiently orthogonal to the columns of A, so it does not absorb signals with meaningful information.
- (ii) Columns of C should be uniformly distributed on the unit sphere  $\mathbb{S}^{N-1}$  so that we could approximate well a typical noise vector.
- (iii) The number of columns of C should grow slower than exponential with N, otherwise the method is impractical.

One way to guarantee all three properties is the *deterministic* approach that consists in filling up C imposing

$$|\langle \boldsymbol{a}_i, \boldsymbol{c}_j \rangle| < rac{lpha}{\sqrt{N}} \; orall i, j \;, \; ext{and} \; |\langle \boldsymbol{c}_i, \boldsymbol{c}_j 
angle| < rac{lpha}{\sqrt{N}} \; orall i 
eq j,$$

with  $\alpha > 1$ . Then it can be shown that the number  $\Sigma$  of columns in C grows at most polynomially :  $N^{\alpha} \leq \Sigma \leq N^{\alpha^2}$ .

We consider instead the *random* approach : the columns of C are drawn at random independently. Then the above inequalities hold but decoherence constraint is weakened by a logarithmic factor  $\sqrt{\ln N}$ .

イロト 不得下 イヨト イヨト

Choose  $\beta > 1$ , and pick  $\Sigma = N^{\beta}$  vectors  $c_i$  at random and independently on  $\mathbb{S}^{N-1}$ .

・ロト ・回ト ・ヨト

Choose  $\beta > 1$ , and pick  $\Sigma = N^{\beta}$  vectors  $c_i$  at random and independently on  $\mathbb{S}^{N-1}$ . Define the convex hulls

$$H_1 = \left\{ x \in \mathbb{R}^N \left| x = \sum_{i=1}^{\Sigma} \xi_i \boldsymbol{c}_i, \ \sum_{i=1}^{\Sigma} |\xi_i| \leqslant 1 \right\}, \quad H_2 = \left\{ x \in \mathbb{R}^N \left| x = \sum_{i=1}^{K} \xi_i \boldsymbol{a}_i, \ \sum_{i=1}^{\Sigma} |\xi_i| \leqslant 1 \right\},\right.$$

・ロト ・回ト ・ヨト

Choose  $\beta > 1$ , and pick  $\Sigma = N^{\beta}$  vectors  $c_i$  at random and independently on  $\mathbb{S}^{N-1}$ . Define the convex hulls

$$H_1 = \left\{ x \in \mathbb{R}^N \left| x = \sum_{i=1}^{\Sigma} \xi_i \boldsymbol{c}_i, \ \sum_{i=1}^{\Sigma} |\xi_i| \leqslant 1 \right\}, \quad H_2 = \left\{ x \in \mathbb{R}^N \left| x = \sum_{i=1}^{K} \xi_i \boldsymbol{a}_i, \ \sum_{i=1}^{\Sigma} |\xi_i| \leqslant 1 \right\},\right.$$

Then, for any  $\kappa > 0$  there exist constants  $\alpha = \sqrt{(\beta - 1)/2}$ ,  $c_0(\kappa, \beta)$  and  $N_0 = N_0(\kappa, \beta)$ , such that  $\forall N > N_0$ 

(i) 
$$\max(\max_{i \leq K} (|\langle \boldsymbol{a}_i, \boldsymbol{\delta b} \rangle|), \max_{i \leq \Sigma} (|\langle \boldsymbol{c}_i, \boldsymbol{\delta b} \rangle|)) < c_0 \frac{\sqrt{\ln N}}{\sqrt{N}},$$

and

(*ii*) 
$$\alpha \frac{\sqrt{\ln N}}{\sqrt{N}} \boldsymbol{\delta b} \in H_1,$$

with probability  $1 - 1/N^{\kappa}$ .

A D F A A F F A

Choose  $\beta > 1$ , and pick  $\Sigma = N^{\beta}$  vectors  $c_i$  at random and independently on  $\mathbb{S}^{N-1}$ . Define the convex hulls

$$H_1 = \left\{ x \in \mathbb{R}^N \left| x = \sum_{i=1}^{\Sigma} \xi_i \boldsymbol{c}_i, \ \sum_{i=1}^{\Sigma} |\xi_i| \leqslant 1 \right\}, \quad H_2 = \left\{ x \in \mathbb{R}^N \left| x = \sum_{i=1}^{K} \xi_i \boldsymbol{a}_i, \ \sum_{i=1}^{\Sigma} |\xi_i| \leqslant 1 \right\},\right\}$$

Then, for any  $\kappa > 0$  there exist constants  $\alpha = \sqrt{(\beta - 1)/2}$ ,  $c_0(\kappa, \beta)$  and  $N_0 = N_0(\kappa, \beta)$ , such that  $\forall N > N_0$ 

(i) 
$$\max(\max_{i \leq K} (|\langle \boldsymbol{a}_i, \boldsymbol{\delta b} \rangle|), \max_{i \leq \Sigma} (|\langle \boldsymbol{c}_i, \boldsymbol{\delta b} \rangle|)) < c_0 \frac{\sqrt{\ln N}}{\sqrt{N}},$$

and

(*ii*) 
$$\alpha \frac{\sqrt{\ln N}}{\sqrt{N}} \boldsymbol{\delta b} \in H_1,$$

with probability  $1 - 1/N^{\kappa}$ .

(i) states that  $H_1$  and  $H_2$  are contained in the  $\ell_2$ -ball of radius  $c_0\sqrt{\ln N}/\sqrt{N}$  except for a few spikes in statistically insignificant directions.



Choose  $\beta > 1$ , and pick  $\Sigma = N^{\beta}$  vectors  $c_i$  at random and independently on  $\mathbb{S}^{N-1}$ . Define the convex hulls

$$H_1 = \left\{ x \in \mathbb{R}^N \left| x = \sum_{i=1}^{\Sigma} \xi_i \boldsymbol{c}_i, \ \sum_{i=1}^{\Sigma} |\xi_i| \leqslant 1 \right\}, \quad H_2 = \left\{ x \in \mathbb{R}^N \left| x = \sum_{i=1}^{K} \xi_i \boldsymbol{a}_i, \ \sum_{i=1}^{\Sigma} |\xi_i| \leqslant 1 \right\},\right.$$

Then, for any  $\kappa > 0$  there exist constants  $\alpha = \sqrt{(\beta - 1)/2}$ ,  $c_0(\kappa, \beta)$  and  $N_0 = N_0(\kappa, \beta)$ , such that  $\forall N > N_0$ 

(i) 
$$\max(\max_{i \leq K} (|\langle \boldsymbol{a}_i, \boldsymbol{\delta b} \rangle|), \max_{i \leq \Sigma} (|\langle \boldsymbol{c}_i, \boldsymbol{\delta b} \rangle|)) < c_0 \frac{\sqrt{\ln N}}{\sqrt{N}},$$

and

(*ii*) 
$$\alpha \frac{\sqrt{\ln N}}{\sqrt{N}} \boldsymbol{\delta b} \in H_1,$$

with probability  $1 - 1/N^{\kappa}$ .

(i) states that  $H_1$  and  $H_2$  are contained in the  $\ell_2$ -ball of radius  $c_0\sqrt{\ln N}/\sqrt{N}$  except for a few spikes in statistically insignificant directions.

(*ii*) states that  $H_1$  contains an  $\ell_2$ -ball of radius  $\alpha \sqrt{\ln N} / \sqrt{N}$  except for a few statistically insignificant directions.

C. Tsogka



#### False Discovery Rate is zero

We have the following results :

#### Theorem (No phantom signal theorem)

Suppose there is no signal :  $\rho = 0$  and  $\delta b/\|\delta b\|_{l_2}$  is uniformly distributed on the unit sphere. Fix  $\beta > 1$ , and draw  $\Sigma = N^{\beta}$  columns for C, independently, from the uniform distribution on  $\mathbb{S}^{N-1}$ . For any  $\kappa > 0$  there are constants  $\tau = \tau(\kappa, \beta)$  and  $N_0 = N_0(\kappa, \beta)$  such that,  $\forall N > N_0$ ,  $\rho_{\tau} = 0$  with probability  $1 - 1/N^{\kappa}$ .

We see no phantom signals when the algorithm is fed with pure noise. The no phantom weight  $\tau$  may be chosen to be the smallest constant such that the no phantom signal theorem holds.

#### False Discovery Rate is zero

We have the following results :

#### Theorem (No phantom signal theorem)

Suppose there is no signal :  $\rho = 0$  and  $\delta b/\|\delta b\|_{l_2}$  is uniformly distributed on the unit sphere. Fix  $\beta > 1$ , and draw  $\Sigma = N^{\beta}$  columns for C, independently, from the uniform distribution on  $\mathbb{S}^{N-1}$ . For any  $\kappa > 0$  there are constants  $\tau = \tau(\kappa, \beta)$  and  $N_0 = N_0(\kappa, \beta)$  such that,  $\forall N > N_0$ ,  $\rho_{\tau} = 0$  with probability  $1 - 1/N^{\kappa}$ .

We see no phantom signals when the algorithm is fed with pure noise. The no phantom weight  $\tau$  may be chosen to be the smallest constant such that the no phantom signal theorem holds.

#### Theorem

Let  $\rho$  be an M-sparse solution of  $\mathcal{A}\rho = \mathbf{b}$ . Assume  $\kappa$ ,  $\beta$ , the Noise Collector, and the noise are the same as in Theorem above. If the columns of  $\mathcal{A}$  are decoherent :  $|\langle \mathbf{a}_i, \mathbf{a}_j \rangle| \leq \frac{1}{3M}$ , then  $supp(\rho_{\tau}) \subseteq supp(\rho)$  for all  $N > N_0$  with probability  $1 - 1/N^{\kappa}$ .

This means that we have zero false discovery rate for any level of noise and with probability that tends to one as the dimension of the data increases to infinity. No false positives !

C. Tsogka

#### Supports of ho and $ho_{ au}$ agree

#### Theorem

Suppose r is the magnitude of smallest non-zero entry of  $\rho$ . If  $\|\delta b\|_{l_2} \leq c_2 \frac{\|b\|_{l_2}^2}{\|\rho\|_{l_1} \sqrt{\ln N}}$ ,  $c_2 = c_2(\kappa, \beta, r, M)$ , then for all  $N > N_0$ ,  $supp(\rho_{\tau}) = supp(\rho)$ , with probability  $1 - 1/N^{\kappa}$ .

Exact support recovery when the noise is not too large.

(ロ) (四) (三) (三)

#### Supports of ho and $ho_{ au}$ agree

#### Theorem

Suppose r is the magnitude of smallest non-zero entry of  $\rho$ . If  $\|\delta b\|_{l_2} \leq c_2 \frac{\|b\|_{l_2}^2}{\|\rho\|_{l_1}} \frac{\sqrt{N}}{\sqrt{\ln N}}$ ,  $c_2 = c_2(\kappa, \beta, r, M)$ , then for all  $N > N_0$ ,  $supp(\rho_{\tau}) = supp(\rho)$ , with probability  $1 - 1/N^{\kappa}$ .

Exact support recovery when the noise is not too large.

Theorem (Exact Recovery)

If there is no noise  $\delta b = 0$ . Then,  $\rho_{\tau} = \rho$  for all  $M < \frac{2\sqrt{N}}{3c_0\tau\sqrt{\ln N}}$  with probability  $1 - 1/N^{\kappa}$ .

Exact recovery in the noise free case.

(日) (四) (日) (日) (日)

To find the minimizer  ${oldsymbol{
ho}}_{ au}, {oldsymbol{\eta}}_{ au}$ ,

$$\begin{aligned} (\boldsymbol{\rho}_{\tau},\boldsymbol{\eta}_{\tau}) &= \arg\min_{\boldsymbol{\rho},\boldsymbol{\eta}} \left( \tau \|\boldsymbol{\rho}\|_{l_{1}} + \|\boldsymbol{\eta}\|_{l_{1}} \right), \\ \text{subject to } \mathcal{A}\boldsymbol{\rho} + \mathcal{C}\boldsymbol{\eta} &= \boldsymbol{b} + \boldsymbol{\delta}\boldsymbol{b}, \end{aligned}$$

we consider a variational approach. We define the function

$$F(x,\eta,\mathrm{z}) = \lambda \left( au \|x\|_{\ell_1} + \|\eta\|_{\ell_1} 
ight) + rac{1}{2} \|\mathcal{A}x + \mathcal{C}\eta - b\|_{\ell_2}^2 + \langle \mathrm{z}, b - \mathcal{A}x - \mathcal{C}\eta 
angle$$

for an O(1) weight  $\tau$ , and determine the solution as

 $\max_{\mathbf{z}} \min_{x,\eta} F(x,\eta,\mathbf{z}).$ 

This variational principle finds the minimum exactly for all values of the regularization parameter  $\lambda$ .

(日) (四) (三) (三)

To determine the exact extremum, we use the iterative soft thresholding algorithm GeLMA which is a semi-implicit version of the primal-dual method of Chambolle & Pock.

- M. Moscoso, A. Novikov, G. Papanicolaou and L. Ryzhik, A differential equations approach to I1-minimization with applications to array imaging, Inverse Problems 28 (2012).
- A. Chambolle and T. Pock, A first-order primal-dual algorithm for convex problems with applications to imaging, Journal of Mathematical Imaging and Vision 40 (2011), pp 120–145.

Pick a value for the regularization parameter  $\lambda$ , e.g.  $\lambda = 1$ . Choose step sizes  $\Delta t_1 < 2/\|[\mathcal{A} | \mathcal{C}]\|^2$  and  $\Delta t_2 < \lambda/\|\mathcal{A}\|$ . Set  $\rho_0 = 0$ ,  $\eta_0 = 0$ ,  $z_0 = 0$ , and iterate for  $k \ge 0$ :

$$egin{aligned} m{r} &= m{b} - \mathcal{A} \, m{
ho}_k - \mathcal{C} \, m{\eta}_k \,, \ m{
ho}_{k+1} &= \mathcal{S}_{ au \, \lambda \Delta t_1} \left( m{
ho}_k + \Delta t_1 \, \mathcal{A}^*(m{z}_k + m{r}) 
ight) \,, \ m{\eta}_{k+1} &= \mathcal{S}_{\lambda \Delta t_1} \left( m{\eta}_k + \Delta t_1 \, \mathcal{C}^*(m{z}_k + m{r}) 
ight) \,, \ m{z}_{k+1} &= m{z}_k + \Delta t_2 \, m{r} \,, \end{aligned}$$

where  $S_{\lambda}(y_i) = \operatorname{sign}(y_i) \max\{0, |y_i| - \lambda\}.$ 

Note : Choosing two step sizes instead of the smaller one  $\Delta t_1$  improves the convergence speed.

• The Noise Collector matrix C is constructed theoretically by drawing  $\Sigma = N^{\beta}$  normally distributed N-dimensional vectors, normalized to unit length.

(日) (日) (日) (日)

- The Noise Collector matrix C is constructed theoretically by drawing  $\Sigma = N^{\beta}$  normally distributed N-dimensional vectors, normalized to unit length.
- In practice we construct C by drawing N<sup>β-1</sup> normally distributed N-dimensional vectors. These are the generating vectors of the Noise Collector. From each one of them a circulant N × N matrix C<sub>i</sub>, i = 1,..., N<sup>β-1</sup>, is constructed. Typically β ≈ 1.5.

- The Noise Collector matrix C is constructed theoretically by drawing  $\Sigma = N^{\beta}$  normally distributed N-dimensional vectors, normalized to unit length.
- In practice we construct C by drawing  $N^{\beta-1}$  normally distributed N-dimensional vectors. These are the generating vectors of the *Noise Collector*. From each one of them a circulant  $N \times N$  matrix  $C_i$ ,  $i = 1, \ldots, N^{\beta-1}$ , is constructed. Typically  $\beta \approx 1.5$ .
- The full Noise Collector matrix can be constructed by concatenation,

 $\mathcal{C} = [\mathcal{C}_1 | \mathcal{C}_2 | \dots | \mathcal{C}_{N^{\beta-1}}]$ 

The full Noise Collector matrix is never formed, only the  $N^{\beta-1}$  generating vectors are stored.

・ロト ・四ト ・ヨト ・ヨト

- The Noise Collector matrix C is constructed theoretically by drawing  $\Sigma = N^{\beta}$  normally distributed N-dimensional vectors, normalized to unit length.
- In practice we construct C by drawing  $N^{\beta-1}$  normally distributed N-dimensional vectors. These are the generating vectors of the *Noise Collector*. From each one of them a circulant  $N \times N$  matrix  $C_i$ ,  $i = 1, \ldots, N^{\beta-1}$ , is constructed. Typically  $\beta \approx 1.5$ .
- The full Noise Collector matrix can be constructed by concatenation,

 $\mathcal{C} = [\mathcal{C}_1 | \mathcal{C}_2 | \dots | \mathcal{C}_{N^{\beta-1}}]$ 

The full Noise Collector matrix is never formed, only the  $N^{\beta-1}$  generating vectors are stored.

• Exploiting the circulant structure of the matrices  $C_i$ , we perform the matrix vector multiplications  $C\eta_k$  and  $C^*(z_k + r)$  using the FFT.

(日) (四) (日) (日) (日)

- The Noise Collector matrix C is constructed theoretically by drawing  $\Sigma = N^{\beta}$  normally distributed N-dimensional vectors, normalized to unit length.
- In practice we construct C by drawing  $N^{\beta-1}$  normally distributed N-dimensional vectors. These are the generating vectors of the *Noise Collector*. From each one of them a circulant  $N \times N$  matrix  $C_i$ ,  $i = 1, \ldots, N^{\beta-1}$ , is constructed. Typically  $\beta \approx 1.5$ .
- The full Noise Collector matrix can be constructed by concatenation,

 $\mathcal{C} = [\mathcal{C}_1 | \mathcal{C}_2 | \dots | \mathcal{C}_{N^{\beta-1}}]$ 

The full Noise Collector matrix is never formed, only the  $N^{\beta-1}$  generating vectors are stored.

- Exploiting the circulant structure of the matrices  $C_i$ , we perform the matrix vector multiplications  $C\eta_k$  and  $C^*(z_k + r)$  using the FFT.
- This makes the complexity associated to the Noise Collector  $O(N^{\beta} \log(N))$ .

< ロ > < 同 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ >

- The Noise Collector matrix C is constructed theoretically by drawing  $\Sigma = N^{\beta}$  normally distributed N-dimensional vectors, normalized to unit length.
- In practice we construct C by drawing  $N^{\beta-1}$  normally distributed N-dimensional vectors. These are the generating vectors of the *Noise Collector*. From each one of them a circulant  $N \times N$  matrix  $C_i$ ,  $i = 1, \ldots, N^{\beta-1}$ , is constructed. Typically  $\beta \approx 1.5$ .
- The full Noise Collector matrix can be constructed by concatenation,

 $\mathcal{C} = [\mathcal{C}_1 | \mathcal{C}_2 | \dots | \mathcal{C}_{N^{\beta-1}}]$ 

The full Noise Collector matrix is never formed, only the  $N^{\beta-1}$  generating vectors are stored.

- Exploiting the circulant structure of the matrices  $C_i$ , we perform the matrix vector multiplications  $C\eta_k$  and  $C^*(z_k + r)$  using the FFT.
- This makes the complexity associated to the *Noise Collector*  $O(N^{\beta} \log(N))$ .
- The cost of performing the matrix vector multiplication  $\mathcal{A} \rho_k$  is NK and, thus, the additional cost due to the Noise Collector is negligible as, typically,  $K \gg N^{\beta-1}$ .

イロト イヨト イヨト イヨト

#### Results



Top row : the images. Bottom row : solution vector with red stars and the true solution vector with green circles.

C. Tsogka	The Noise Collector	October 1, 2020	23 / 29
		・ロト ・四ト ・ヨト ・ヨト 三百	$\mathcal{O}\mathcal{A}\mathcal{O}$

#### Results



Top row : the images. Bottom row : solution vector with red stars and the true solution vector with green circles.

C. Tsogka	The Noise Collector	October 1, 2020	23 / 29
		・ロト ・四ト ・ヨト ・ヨト 三日	$\mathcal{O}\mathcal{A}\mathcal{O}$

#### Results



Top row : the images. Bottom row : solution vector with red stars and the true solution vector with green circles.

$$\beta = 1.5, \ SNR = \frac{\|\boldsymbol{b}\|_{l_2}}{\|\boldsymbol{\delta}\boldsymbol{b}\|_{l_2}} = 1, \ N = 625, \ K = 1681, \ \tau = 2.$$

### The Noise Collector at work



### The Noise Collector at work



#### Failure to recover



abscissa are the sparsity M and  $\frac{\|\delta b\|_2}{\|b\|_2}$ . The black line is  $\frac{1}{\sqrt{\ln N}} \frac{\sqrt{N}}{\sqrt{M}}$ .

### Table of contents

Inverse problems in wave propagation

2 Applications

(3) On  $\ell_1$  and (super)resolution

4 Uncertainty in the data

#### 5 Conclusion

・ロト ・回ト ・ヨト

### Concluding remarks

• Provided a theoretical framework to examine under what conditions the  $\ell_1$ -minimization problem admits a solution that is close to the exact one when coherence in the columns of  $\mathcal{A}$  is allowed inside the vicinities.

A D F A A F F A

- Provided a theoretical framework to examine under what conditions the  $\ell_1$ -minimization problem admits a solution that is close to the exact one when coherence in the columns of  $\mathcal{A}$  is allowed inside the vicinities.
- To increase the robustness of  $\ell_1$ -minimization when imaging with noisy data we propose to solve

 $\begin{aligned} (\boldsymbol{\rho}_{\tau}, \boldsymbol{\eta}_{\tau}) &= \arg\min_{\boldsymbol{\rho}, \boldsymbol{\eta}} \left( \tau \|\boldsymbol{\rho}\|_{l_{1}} + \|\boldsymbol{\eta}\|_{l_{1}} \right), \\ \text{subject to } \mathcal{A}\boldsymbol{\rho} + \mathcal{C}\boldsymbol{\eta} &= \boldsymbol{b} + \boldsymbol{\delta}\boldsymbol{b}, \end{aligned}$ 

instead of

$$oldsymbol{
ho}_* = rg\min_{oldsymbol{
ho}} \|oldsymbol{
ho}\|_{\ell_1}, \ {
m subject to} \ {\cal A} \ oldsymbol{
ho} = oldsymbol{b} + \delta oldsymbol{b},$$

Here the  $N \times N^{\beta}$  matrix C is the *Noise Collector* matrix and  $\tau$  is an O(1) weight for *any* level of noise.

- Provided a theoretical framework to examine under what conditions the  $\ell_1$ -minimization problem admits a solution that is close to the exact one when coherence in the columns of  $\mathcal{A}$  is allowed inside the vicinities.
- To increase the robustness of  $\ell_1$ -minimization when imaging with noisy data we propose to solve

 $\begin{aligned} (\boldsymbol{\rho}_{\tau}, \boldsymbol{\eta}_{\tau}) &= \arg\min_{\boldsymbol{\rho}, \boldsymbol{\eta}} \left( \tau \|\boldsymbol{\rho}\|_{l_{1}} + \|\boldsymbol{\eta}\|_{l_{1}} \right), \\ \text{subject to } \mathcal{A}\boldsymbol{\rho} + \mathcal{C}\boldsymbol{\eta} &= \boldsymbol{b} + \delta\boldsymbol{b}, \end{aligned}$ 

instead of

$$\rho_* = \arg\min_{\rho} \|\rho\|_{\ell_1}, \text{ subject to } \mathcal{A} \rho = b + \delta b,$$

Here the  $N \times N^{\beta}$  matrix C is the *Noise Collector* matrix and  $\tau$  is an O(1) weight for *any* level of noise.

• The additional cost of the Noise Collector is negligible compared to the cost of solving the original problem.

(日) (四) (日) (日) (日)

## Concluding remarks

• The most important features of the Noise Collector are that :

・ロト ・回ト ・ヨト ・

## Concluding remarks

- The most important features of the Noise Collector are that :
  - In calibration is necessary with respect to the level of noise,

・ロト ・回ト ・ヨト

#### Conclusion

### Concluding remarks

- The most important features of the Noise Collector are that :
  - In calibration is necessary with respect to the level of noise,
  - 2 exact support recovery is obtained for relatively large levels of noise

 $\|\boldsymbol{\delta}b\|_{\ell_2}\leqslant c_1rac{\|\boldsymbol{b}\|_{\ell_2}^2\sqrt{N}}{\|\boldsymbol{\rho}\|_{\ell_1}\sqrt{\ln N}}$ ,

イロト イロト イヨト

#### Conclusion

### Concluding remarks

- The most important features of the Noise Collector are that :
  - In calibration is necessary with respect to the level of noise,
  - 2 exact support recovery is obtained for relatively large levels of noise

 $\|\boldsymbol{\delta}b\|_{\ell_2} \leqslant c_1 \frac{\|\boldsymbol{b}\|_{\ell_2}^2 \sqrt{N}}{\|\boldsymbol{\rho}\|_{\ell_1} \sqrt{\ln N}},$ 

9 we have zero false discovery rates for all levels of noise, with high probability.

#### Conclusion

### Concluding remarks

- The most important features of the Noise Collector are that :
  - In calibration is necessary with respect to the level of noise,
  - (a) exact support recovery is obtained for relatively large levels of noise

 $\|\boldsymbol{\delta}b\|_{\ell_2}\leqslant c_1rac{\|\boldsymbol{b}\|_{\ell_2}^2\sqrt{N}}{\|\boldsymbol{
ho}\|_{\ell_1}\sqrt{\ln N}}$ ,

9 we have zero false discovery rates for all levels of noise, with high probability.

- More on the Noise Collector and its theoretical analysis in
  - M. Moscoso, A. Novikov, G. Papanicolaou, CT, *Imaging with highly incomplete and corrupted data*, Inverse Problems, 36(3), p. 035010, 2020. https://doi.org/10.1088/1361-6420/ab5a21
  - M. Moscoso, A. Novikov, G. Papanicolaou, CT, The Noise Collector for sparse recovery in high dimensions, Proceedings of the National Academy of Sciences, 117 (21), p. 11226-11232, 2020. https://doi.org/10.1073/pnas.1913995117

(日) (同) (三) (三)